# Analysing satellite data for IO vertical columns in polar and tropical regions

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Nucleation procedure

# 1) Motivation – iodine species in the troposphere

Why is iodine important for tropospheric composition?

- Strong ozone depletion potential via catalytic cycles
- Change of oxidation pathways
- Nucleation of higher iodine oxides  $I_xO_v$  (e.g.  $I_2O_5$ ,  $I_2O_4$ )
- Possible growth to cloud condensation nuclei  $\rightarrow$  Impact on radiation balance





### Sources of atmospheric iodine

- Mainly maritime sources identified, release pathways not yet fully understood
- Biogenic release by certain types of algae/phytoplankton: I<sub>2</sub>, CH<sub>2</sub>I<sub>2</sub>, CHICI, etc
- Inorganic release: e.g. surface reactions of  $O_3$  with  $I^2$ , yet unknown pathways

## 2) The SCIAMACHY instrument

SCanning Imaging Absorption spectroMeter for Atmospheric CHartographY

- UV-Vis-NIR spectrometer onboard ENVISAT
- spectral range within 214 2400 nm
- sun-synchronous orbit at 800 km altitude
- observation geometries nadir, limb, occultation
- ground pixel size typically 30 x 60 km<sup>2</sup>
- launch in 2002, mission assured until 2014
- mission might be further extended

### 3) The IO retrieval by DOAS

### **DOAS retrieval settings for IO**

Fitting window: 416 to 430 nm (2 absorption bands) NO<sub>2</sub> (223K), O<sub>3</sub> (221K), IO (298K) Trace gases: Ring effect, stray light, 2<sup>nd</sup> ord. polynomial Other features: Slant column amounts (SC) of the trace gases Result:

**Relevant definitions** 

SC (slant column): Trace gas amount integrated along the individual light path and then averaged over all occurring light paths.

VC (vertical column): Trace gas amount integrated over all altitudes vertically above a ground area of 1 cm<sup>2</sup>.

**AMF (air mass factor):** Light path enhancement/reduction within the absorber layer, i.e. also equivalent to the ratio between SC and VC.

**BAMF (block air mass factor)**: Discrete change in the retrieved quantity (here SC), if the actual vertical column changes by a discrete amount at a certain altitude interval of the discrete altitude grid, hence representing the altitude sensitivity of the retrieval.





SCIAMACHY onboard ENVISAT, Monitoring the Changing Earth's Atmosphere, published by DLR, 2006. (ESA, artist's impression)

> **D**ifferential **O**ptical Absorption **S**pectroscopy

### 4) AMF considerations



### Applied code: SCIATRAN (Rozanov, et al. 2005)

Figure 1: Block AMF values for IO calculated for a Rayleigh atmosphere (no aerosol) and for different cases of surface albedo. SZA: 70°

	()	NOZUNC
Air Mass Factor	5.0	— Laver Height 1km
	4.8	- Layer Height 100
	4.5	
	4.3	-
	4.0	
	3.8	
	3.5	
	3.3	
	3.0	30 40



McFiggans et al, 2004.

### 5) IO vertical columns above Antarctica





- Radiative transfer calculations show that the satellite sensitivity for IO is largest above bright surfaces such as snow and ice. In this case, the AMF depends only weakly on the IO profile.
- Six years of IO observations show many details of spatial and temporal variation above th Antarctic.
- •Occurence of IO above sea ice in late spring may be linked to biogenic emissions from below the ice.
- •Observations of IO above the tropical oceans are subject to larger noise and uncertainties than over snow and ice.
- Enhancements of tropical IO might be linked to biological activity in upwelling regions of the ocean.

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Figure 2: AMF values for IO with respect to varying SZA at 90% albedo for two different heights: box profile volume constant mixing ratio (VMR) up to 1km (black) and up to 100m (grey).



Figure 3: AMF values for IO with respect to varying albedo for box profile heights (BPH) of 1km (blue), 500m (turqouise, used in Fig. 6) and 100m (light blue).

Selected References

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>The lower the albedo, the lower the sensitivity for IO at the surface, i.e. IO above snow/ice is better visible than over ocean.

>Over bright surfaces, the IO box profile height (e.g. 100m or 1km) has no major influence on the AMF, differences remain below 7%. Over dark surfaces, typical deviations are 20-40%.

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see also: www.iup.physik.uni-bremen.de/doas